# Emission spectra of atoms with non-Markovian interaction: Fluorescence in a photonic crystal 

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#### Abstract

We present a formula to evaluate the spontaneous emission spectra of an atom in contact with a radiation field with non-Markovian effects. This formula is written in terms of a two-time correlation of system observables and the environmental correlation function, and depends on the distance between the emitting atom and the detector. As an example, we use it to analyze the fluorescence spectra of a two level atom placed as an impurity in a photonic crystal. The radiation field within those materials has a gap or discontinuity where electromagnetic modes cannot propagate in the stationary limit. In that situation, the atomic emission occurs in the form of evanescent waves which are detected with less efficiency the farther we place the detector. The methodology presented in this paper may be useful to study the non-Markovian dynamics of any quantum open system in linear interaction with a harmonic oscillator reservoir and within the weak-coupling approximation.


DOI: 10.1103/PhysRevA.77.043836

## I. INTRODUCTION

Most of the theoretical work developed in the context of quantum open systems relies in the use of the Markov approximation, which assumes that when a quantum open system (e.g., an atom) interacts with its environment (e.g., a radiation field), the later is not affected in a time scale comparable to that of the system.

However, the Markovian assumption is only valid for radiation fields in which the photonic density of states (DOS) is a continuous and smoothly varying function in the frequency space. When this function varies in a range which is comparable to the atomic spontaneous emission rate $\Gamma$, then the environmental relaxation time (also correlation time) $\tau_{c}$, can no longer be considered negligible in comparison with the atomic relaxation time. We are then dealing with nonMarkovian interactions, in which the environment does not absorb instantaneously every photon it receives from the atom, but needs a certain memory time $\tau_{c}$ to do it and return to its equilibrium state.

Non-Markovian interactions may occur when the radiation field is placed in some structured environment, so that its photonic DOS may present sudden jumps or singular behavior in a frequency range comparable to the spontaneous emission rate. For instance, in the so-called photonic crystals (PCs) [1,2], there is a periodicity in the refraction index which produces the scattering of light outside of the crystal for certain frequencies that are related to the lattice periodicity. Since these modes are absent from the crystal, the photonic DOS is zero in the corresponding frequencies, so that a gap is formed. The frequency ranges in which the photonic DOS is nonzero are known as bands. Within the bands, the photonic DOS varies in frequency range comparable to the atomic spontaneous emission rate, which gives rise to a long memory time for the environment. This kind of nonMarkovian interaction may also be found in a quantum cavity, when the range of variation of the photonic DOS is
comparable to the spontaneous emission rate [3-5].
In non-Markovian interactions the emitted photon may be eventually reabsorbed by the atom, which produces a backaction in the system dynamics, both in the evolution of its expectation values and the multiple-time correlation functions (MTCFs) of its observables. An important consequence of this is that some related physical quantities such as the atomic emission spectra also display non-Markovian features. Hence, it is necessary to obtain a more general formula for the atomic emission spectra without using the Markovian approximation.

In most of the literature (see, for instance, Ref. [3], and references therein), the analytical expression for the spectra is derived by assuming the Markov approximation at some stage. To be more specific, in photodetector experiments the emission spectra depends on a two-time correlation of the positive and negative operators of the radiation field emitted by the atom. In order to evaluate it, this two-time correlation is expressed in terms of a two-time correlation of system observables, which is done by assuming the Markovian approximation. In the end, it is found that the spectra is related to a Fourier transform of the system two-time correlation [3,6]. Moreover, once the two-time correlation is computed through the quantum regression theorem, a Lorentzian profile for the spectra is obtained. This theorem states that the evolution of N -time correlation functions can be computed with the master equation by only considering a rotation in its initial condition.

In this paper, we derive a formula to compute the atomic emission spectra for non-Markovian interactions. We show that the two-time correlation of the positive and negative operators of the radiation field is related to a two-time correlation of system observables through a double convolution integral that also depends on the environmental correlation function. The decaying of this environmental correlation function gives the correlation time or relaxation time of the environment $\tau_{c}$, which for non-Markovian interactions is nonzero. Moreover, the two-time correlation of the system
observables can no longer be described within the so-called quantum regression theorem (QRT), since the derivation of such formula relies in the Markov approximation [7-9]. In the non-Markovian case, the set of evolution equations derived in Refs. [10,11] should be used to evaluate MTCF. Moreover, the formula we derive for the emission spectra is dependent on the spatial distance between the emitting atom and the detector.

In order to illustrate the theory, we study the fluorescence spectra of a two-level atom in contact with the modified radiation field that exists within a photonic crystal. In the stationary limit, the emission in the frequency region corresponding to the gap is made in the form of evanescent modes, which present an exponentially decaying spatial dependency. It is clear that when evanescent modes are present in the emission, the fluorescence spectra depends dramatically on the distance from the atom to the detector.

The plan of the paper is the following. In Sec. II we present the general Hamiltonian for the class of systems that we want to describe, as well as the environmental correlation function that characterizes the interaction. In Sec. III we present a formula to obtain a non-Markovian spectra in terms of a two-time correlation of system observables, convoluted with the environmental correlation function. This formula depends explicitly on the distance between the emitting system and the detector, and is one of the main results of this paper. Section IV is devoted to review the evolution equations of system two-time correlations within the weak coupling limit. In the last part of the paper (Sec. V) we apply the theory to compute the fluorescence spectra of a two-level atom placed as an impurity in a photonic crystal, showing that the distance between the atom and the detector is an important parameter to take into account if we want to characterize correctly the spectra in the gap region.

## II. NON-MARKOVIAN INTERACTIONS

Let us consider the model Hamiltonian to study the dynamics of $\mathcal{S}$ with Hamiltonian $H_{S}$, in interaction with $H_{B}$. Assuming a linear coupling, we have

$$
\begin{equation*}
H_{\mathrm{tot}}=H_{S}+H_{B}+H_{I}=H_{S}+\sum_{\lambda} \omega_{\lambda} a_{\lambda}^{\dagger} a_{\lambda}+\sum_{\lambda} g_{\lambda}\left(a_{\lambda}^{\dagger} L+L^{\dagger} a_{\lambda}\right), \tag{1}
\end{equation*}
$$

where the operator $L$ acts on the Hilbert space of the system and $a_{\lambda}, a_{\lambda}^{\dagger}$ are the annihilation and creation operators on the environment Hilbert space. The $g_{\lambda}^{\prime}$ 's are the coupling constants that can be taken as real numbers, and the $\omega_{\lambda}^{\prime}$ 's are the frequencies of the harmonic oscillators that constitute the environment.

Along the paper we will assume the weak coupling approximation, which consists in considering that the system interaction Hamiltonian is $g$ times smaller than the free Hamiltonian, which includes both the Hamiltonian of the system and the environment. We will also use for simplicity units in which $\hbar=1$.

The Hamiltonian (1) is good enough to characterize most of the interactions in quantum optics related to processes in
which only one photon is involved, and which can be described within the dipolar approximation. The dynamics of the quantum mean values and MTCF can be computed in two distinct ways. First, by evolving vectors on the system Hilbert space (see, for instance, Refs. [10,11]). In this evolution there are two important quantities that describe the action of the environment into the system: the environmental correlation function

$$
\begin{equation*}
\alpha(t-\tau)=\sum_{\lambda} g_{\lambda}^{2} e^{-i \omega_{\lambda}(t-\tau)}, \tag{2}
\end{equation*}
$$

which as noted before decays within a time $\tau_{c}$ and describes the dissipation of the system, and a certain quantity $z_{t}$, such that $\alpha(t-\tau)=\mathcal{M}\left[z_{t} z_{\tau}^{*}\right]$, where $\mathcal{M}[\cdots]$ represents a Gaussian average. The function $z_{t}$ can be interpreted as a Gaussian colored noise, and describes the stochastic effect of the many environmental degrees on the system wave vector [12-16]. Second, in the Heisenberg picture, by evolving system operators. In these equations, the action of the environment is described entirely through the correlation function (2). Note that in the Markov case, $\alpha(t-\tau) \approx \delta(t-\tau)$ (corresponding to $\tau_{c} \approx 0$ ), and the quantity $z_{t}$ may represent a Gaussian white noise.

In this paper, we need to compute some two-time correlations of system observables. We will do it within the Heisenberg picture scheme, by using the system of equations derived in Refs. [10,11] in the weak-coupling limit.

To compute the environment correlation function, it is necessary to perform the sum appearing in Eq. (2), which can, in principle, be done once the value of $g_{\lambda}$ is known for the particular interaction. For instance, in the our case of an atom dipolarly coupled with the radiation field, the coupling constants are known, and the quantum number is $\lambda \equiv\{\mathbf{k}, \sigma\}$, where $\mathbf{k}$ is the wave vector and $\sigma$ corresponds to the two polarization modes.

However, there are certain problems where the functional form $g_{\lambda}$ is not known, for instance atoms in contact with the phonons in a solid. In those cases, it is often useful to express the sum appearing in Eq. (2) as

$$
\begin{equation*}
\alpha(t-\tau)=\int d \omega J(\omega) e^{-i \omega(t-\tau)} \tag{3}
\end{equation*}
$$

where we have defined the so-called spectral function as $J(\omega)=\pi \Sigma_{\lambda} g_{\lambda}^{2} \delta\left(\omega-\omega_{\lambda}\right)$. This spectral function can be approximated with some phenomenological models (see, for instance, Refs. [17,18]). The spectral function can be expressed as $J(\omega)=G^{2}(\omega) \rho(\omega)$, where $G(\omega)$ is the function $g_{\lambda}$ in the continuum and $\rho(\omega)$ is the photonic DOS of the environment. Inserting this relation in Eq. (3) it is easy to see the important role of the photonic DOS in the correlation function, and particularly in its decaying time $\tau_{c}$.

Finally, we note that despite the weak coupling approximation here used is only valid for weak and medium couplings, and therefore it might not capture completely the non-Markovian behavior of the system, it does describe some important memory effects in its dynamics. Particularly,
the evolution of the system observables depend on their past history up to the time $\tau_{c}$ that the environment takes to relax back to equilibrium after the interaction.

## III. ATOMIC EMISSION SPECTRA

In this section we present the formula necessary to obtain the emission spectra of an atom with non-Markovian interaction. To this end, we follow a well-known photodetection model of experiment, the gedanken spectrum analyzer that can be found in Refs. [3,19], and which provides an operational definition of the spectral profile. In this scheme, we consider that the radiation field emitted by a two-level atom is detected by another two-level atom with frequency $\omega$, which is initially prepared in its ground state $|1\rangle$. The detecting atom has Hamiltonian $H_{\mathcal{D}}=\omega \sigma_{z} / 2$, and it is placed inside a shutter, that only opens during a certain observation time $T$, during which it receives the emitted radiation, being eventually excited. The Hamiltonian of the emitting atom (with levels $|1\rangle$ and $|2\rangle$ ) is given by

$$
\begin{equation*}
H_{S}=-\frac{\omega_{12}}{2}\left(\sigma_{22}-\sigma_{11}\right)=\frac{\omega_{12}}{2} \sigma_{z}, \tag{4}
\end{equation*}
$$

and the total Hamiltonian of emitting atom and radiation field is described by a Hamiltonian $H_{R}$, given by $H_{R}=H_{S}$ $+H_{B}+\Sigma_{\lambda} g_{\lambda}\left(L^{\dagger} a_{\lambda}+a_{\lambda}^{\dagger} L\right)$.

The Hamiltonian of the total system (detector atom, emitting atom, and radiation field) reads as follows:

$$
\begin{equation*}
H=H_{\mathcal{D}}+H_{R}+W \tag{5}
\end{equation*}
$$

Here the coupling between the detecting atom $H_{\mathcal{D}}$ with the radiation field, with Hamiltonian $H_{R}$, is dipolar and given by a Hamiltonian $W$, which in interaction image with respect to the detector is given by

$$
\begin{equation*}
\widetilde{W}(t)=\left[\sigma_{21} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}(\mathbf{r}, t) e^{i \omega t}+\sigma_{12} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}(\mathbf{r}, t) e^{-i \omega t}\right] \tag{6}
\end{equation*}
$$

where we have considered $d_{21}^{\mathcal{D}} \hat{\mathbf{d}}^{\mathcal{D}}=d_{12}^{\mathcal{D}} \hat{\mathbf{d}}^{\mathcal{D}}=\langle 1| \mathbf{D}|2\rangle=\mathbf{d}^{\mathcal{D}}$. The superindex $\mathcal{D}$ reminds that these are the components of the detector's dipole, and $\mathbf{D}$ is the dipolar operator. It is important to note here that the field operators $\mathbf{E}^{(+)}$and $\mathbf{E}^{(-)}$, correspond to the radiative atoms and the radiation field with $H_{R}$. The positive part of the field placed in $\mathbf{r}$ is defined as

$$
\begin{equation*}
\mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)=\sum_{\lambda} \epsilon_{\lambda} A_{\lambda}(\mathbf{r}) a_{\lambda}\left(\mathbf{r}_{0}, t\right) \mathbf{e}_{\lambda} \tag{7}
\end{equation*}
$$

and $\mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)=\left[\mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)\right]^{\dagger}[19]$. In the last expression (and from now on) we have added explicitly the dependency on the position $\mathbf{r}_{0}$ of the source dipole that originates the field. The quantity $\epsilon_{\lambda}=\sqrt{\frac{\omega_{\lambda}}{2 \epsilon_{0} v}}$, with $v$ the quantization volume. In terms of the coupling strengths we find that $g_{\lambda}$ $\equiv g_{\lambda}(\mathbf{r})=\epsilon_{\lambda} A_{\lambda}(\mathbf{r}) \mathbf{d} \cdot \mathbf{e}_{\lambda}$.

As a condition, the time of observation $T$ is much bigger that the inverse of the natural width $\Gamma$ of the detecting atom excited level. In addition, the Rabi frequency of the emitting atom has to be larger than the inverse of $T$. With this set up, the idea is to calculate the spectral distribution of the fluorescence light $P(\omega, T)$. This is defined as the probability of
excitation of the detecting atom at the time of observation $T$, i.e.,

$$
\begin{equation*}
P(\omega, T)=\operatorname{Tr}_{R, \mathcal{D}}[|2\rangle\langle 2| \rho(T)], \tag{8}
\end{equation*}
$$

where $\rho(T)$ is the density matrix of the total system at time $T$. In the interaction representation, this density matrix is,

$$
\begin{align*}
\rho(T)= & \rho(0)-i \int_{0}^{T}[\tilde{W}(t), \rho(0)] \\
& -\int_{0}^{T} d t \int_{0}^{t} d t^{\prime}\left[\tilde{W}(t),\left[\tilde{W}\left(t^{\prime}\right), \rho(0)\right]\right] \tag{9}
\end{align*}
$$

for an expansion in which $\rho(T) \approx \rho(0)$.
Replacing Eq. (9) in Eq. (8), we get the following expression for $P(\omega, T)$ :

$$
\begin{align*}
P(\omega, T)= & \operatorname{Tr}_{R}\left[\int_{0}^{T} d t \int_{0}^{t} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right) \rho_{R}\right. \\
& \times \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t^{\prime}\right)+\int_{0}^{T} d t \int_{0}^{t} d t^{\prime} e^{i \omega\left(t^{\prime}-t\right)} \\
& \left.\times \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t^{\prime}\right) \rho_{R} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)\right] \tag{10}
\end{align*}
$$

If we now change $t$ into $t^{\prime}$, and $t^{\prime}$ into $t$ in the second term,

$$
\begin{align*}
P(\omega, T)= & \operatorname{Tr}_{R}\left[\int_{0}^{T} d t \int_{0}^{t} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right) \rho_{R}\right. \\
& \times \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t^{\prime}\right)+\int_{0}^{T} d t^{\prime} \int_{0}^{t^{\prime}} d t e^{i \omega\left(t-t^{\prime}\right)} \\
& \left.\times \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right) \rho_{R} \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t^{\prime}\right)\right] \tag{11}
\end{align*}
$$

both integrands becomes the same. On the other hand, the regions of integration are the complementary parts of a square of side $T$, so that the last expression can be written as

$$
\begin{equation*}
P(\omega, T)=\int_{0}^{T} d t \int_{0}^{T} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)} g^{(1)}\left(\mathbf{r}, \mathbf{r}_{0} ; t, t^{\prime}\right), \tag{12}
\end{equation*}
$$

where the average $\langle\cdots\rangle=\operatorname{Tr}_{R}\left(\rho_{R} \cdots\right)$, and we have defined

$$
\begin{equation*}
g^{(1)}\left(\mathbf{r}, \mathbf{r}_{0} ; t, t^{\prime}\right)=\left\langle\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right) \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t^{\prime}\right)\right\rangle \tag{13}
\end{equation*}
$$

as the first order correlation of the projection of the emitted field in the direction of the dipole.

In order to express the last equation in terms of observables of the emitting atoms, it is necessary to express the field operator, which is defined in Eq. (7) in terms of the atomic dipolar momentum. We insert in Eq. (7) the solution of the Heisenberg equation for $a_{\lambda}\left(\mathbf{r}_{0}, t\right)$ with the total Hamiltonian (5), i.e., $d a_{\lambda}\left(\mathbf{r}_{0}, t\right) / d t=-i\left[H, a_{\lambda}\left(\mathbf{r}_{0}, t\right)\right]$,

$$
\begin{equation*}
a_{n}\left(\mathbf{r}_{0}, t\right)=a_{\lambda}\left(\mathbf{r}_{0}, 0\right) e^{-i \omega_{\lambda} \tau}-i \int_{0}^{t} d \tau g_{\lambda} L\left(t^{\prime}\right) e^{-i \omega_{\lambda}(t-\tau)} \tag{14}
\end{equation*}
$$

Here, the coupling constant between emitting atoms and surrounding field is $g_{\lambda} \equiv g_{\lambda}\left(\mathbf{r}_{0}\right)=\epsilon_{\lambda} A_{\lambda}\left(\mathbf{r}_{0}\right) \mathbf{d} \cdot \mathbf{e}_{\lambda}$. After this replacement, the component of $\mathbf{E}^{(+)}$in the direction and position of the detector is

$$
\begin{equation*}
\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)=-i \sum_{\lambda} g_{\lambda} g_{\lambda}^{\mathcal{D}} \int_{0}^{t} d \tau L(\tau) e^{-i \omega_{\lambda}(t-\tau)} \tag{15}
\end{equation*}
$$

where we have defined $g_{\lambda}^{\mathcal{D}} \equiv g_{\lambda}^{\mathcal{D}}(\mathbf{r})=\epsilon_{\lambda} A_{\lambda}(\mathbf{r}) \mathbf{d}^{\mathcal{D}} \cdot \mathbf{e}_{\lambda}$ as the coupling constant of the detector with the emitted field. In expression (15) we only keep the part of the field corresponding to the radiated field from the dipole (also "source field"). The other part, corresponding to the quantum vacuum field

$$
\begin{equation*}
\mathbf{E}_{0}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)=\sum_{\lambda} \epsilon_{\lambda} A_{\lambda}(\mathbf{r}) a_{\lambda}\left(\mathbf{r}_{0}, 0\right) e^{-i \omega_{\lambda} t} \mathbf{e}_{\lambda} \tag{16}
\end{equation*}
$$

does not contribute to photodetection signals, since provided that the field is in the vacuum state $|0\rangle$,

$$
\begin{equation*}
\langle 0|\left[\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}_{0}^{(-)}\left(\mathbf{r}, \mathbf{r}_{0}, \tau\right)\right]=0, \quad\left[\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}_{0}^{(+)}\left(\mathbf{r},, \mathbf{r}_{0}, 0\right)\right]|0\rangle=0 \tag{17}
\end{equation*}
$$

in Eq. (13). We now introduce a distribution function $\rho\left(\mathbf{r}_{0}\right)$, which describes the density of radiating atoms at a given positions of the crystal. Assuming that the atomic distribution is the same for each Wigner Seitz cell (WSC) of the crystal [20], we perform an average over the atomic distribution within the crystal, obtaining the following expression for the positive part of the radiation field:

$$
\begin{equation*}
\left\langle\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)\right\rangle_{\mathbf{r}_{0}}=\frac{1}{N_{e}} \int_{\mathrm{WSC}} d \mathbf{r}_{0} \rho\left(\mathbf{r}_{0}\right) \mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right), \tag{18}
\end{equation*}
$$

where $N_{e}=\int_{\mathrm{WSC}} d \mathbf{r}_{0} \rho\left(\mathbf{r}_{0}\right)$ is the number of active atoms in a cell. Considering one single atom at position $\mathbf{r}_{a}$ in each cell, such that $\rho\left(\mathbf{r}_{0}\right)=\delta\left(\mathbf{r}_{0}-\mathbf{r}_{a}\right)$ and $N_{e}=1$. Then we have

$$
\begin{equation*}
\left\langle\hat{\mathbf{d}}_{D} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{0}, t\right)\right\rangle_{\mathbf{r}_{0}}=-i \int_{0}^{t} d \tau \mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right) L(\tau) \tag{19}
\end{equation*}
$$

with

$$
\begin{equation*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right)=\sum_{\lambda} g_{\lambda}^{\mathcal{D}} g_{\lambda} e^{-i \omega_{\lambda}(t-\tau)} e^{i\left(\mathbf{r}-\mathbf{r}_{a}\right) \mathbf{k}} \tag{20}
\end{equation*}
$$

where for simplicity in the notation we set $g_{\lambda} \equiv g_{\lambda}\left(\mathbf{r}_{a}\right)$. Replacing Eq. (19) in Eq. (12), we get

$$
\begin{align*}
g^{(1)}\left(\mathbf{r}, \mathbf{r}_{a}, t, t^{\prime}\right)= & \left\langle g^{(1)}\left(\mathbf{r}, \mathbf{r}_{0}, t, t^{\prime}\right)\right\rangle_{\mathbf{r}_{0}}=\left\langle\left(\left\langle\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(-)}\left(\mathbf{r}, \mathbf{r}_{a}, t\right)\right\rangle_{\mathbf{r}_{0}}\right)\right. \\
& \left.\times\left(\left\langle\mathbf{d}^{\mathcal{D}} \cdot \mathbf{E}^{(+)}\left(\mathbf{r}, \mathbf{r}_{a}, t^{\prime}\right)\right\rangle_{\mathbf{r}_{0}}\right)\right\rangle \\
= & \int_{0}^{t} d \tau \int_{0}^{t^{\prime}} d \tau^{\prime} \mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right) \mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t^{\prime}, \tau^{\prime}\right) \\
& \times\left\langle L^{\dagger}(\tau) L\left(\tau^{\prime}\right)\right\rangle . \tag{21}
\end{align*}
$$

Following Eq. (12) the spectra has the following form:

$$
\begin{align*}
P(\omega, T)= & \int_{0}^{T} d t \int_{0}^{T} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)} g^{(1)}\left(\mathbf{r}, \mathbf{r}_{a}, t, t^{\prime}\right) \\
= & \int_{0}^{T} d t \int_{0}^{T} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)} \\
& \times\left\{\int_{0}^{t} d \tau \int_{0}^{t^{\prime}} d \tau^{\prime} \mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right) \mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t^{\prime}, \tau^{\prime}\right)\right. \\
& \left.\times\left\langle L^{\dagger}(\tau) L\left(\tau^{\prime}\right)\right\rangle\right\} \tag{22}
\end{align*}
$$

In order to compute the fluorescence spectra, the limit of $T$ $\rightarrow \infty$ has to be taken, so that the signal is observed in the stationary limit. In that case, the above formula corresponds to two Laplace transform of a convolution

$$
\begin{equation*}
P(\omega)=\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a},-\omega\right) \mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)\left\langle L^{\dagger}(-\omega) L(\omega)\right\rangle \tag{23}
\end{equation*}
$$

Since

$$
\begin{equation*}
P(\omega, \infty)=P(\omega)=\operatorname{Tr}_{R, \mathcal{D}}[|2\rangle\langle 2| \rho(\infty)] \tag{24}
\end{equation*}
$$

we are dealing with the detector reduced density matrix in the stationary limit.

## No spatial dependency considered

When there is no spatial dependency of the functions $\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right)$, i.e., $e^{i \mathbf{k} \cdot\left(\mathbf{r}_{0}-\mathbf{r}_{a}\right)} \approx 1$ then

$$
\begin{equation*}
\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right)=\alpha(t-\tau) \tag{25}
\end{equation*}
$$

is the usual correlation function $\alpha(t-\tau)$. In that case, we can write Eq. (22) as

$$
\begin{align*}
P(\omega, T)= & \int_{0}^{T} d t \int_{0}^{T} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)}\left\{\int _ { 0 } ^ { t } d \tau \int _ { 0 } ^ { t ^ { \prime } } d \tau ^ { \prime } \alpha ^ { * } ( t - \tau ) \alpha \left(t^{\prime}\right.\right. \\
& \left.\left.-\tau^{\prime}\right)\left\langle L^{\dagger}(\tau) L\left(\tau^{\prime}\right)\right\rangle\right\} \tag{26}
\end{align*}
$$

Hence, the fluorescence spectra can be written as

$$
\begin{equation*}
P(\omega)=\alpha^{*}(-\omega) \alpha(\omega)\left\langle L^{\dagger}(-\omega) L(\omega)\right\rangle \tag{27}
\end{equation*}
$$

In addition, in the Markov case, such correlation is a delta function $\alpha(t-\tau)=\Gamma \delta(t-\tau)$ and the last formula is just

$$
\begin{equation*}
P(\omega, T)=\Gamma^{2} \int_{0}^{T} d t \int_{0}^{T} d t^{\prime} e^{i \omega\left(t-t^{\prime}\right)}\left\langle L^{\dagger}(t) L\left(t^{\prime}\right)\right\rangle, \tag{28}
\end{equation*}
$$

which in the equilibrium, and with the observation time $T$ $\rightarrow \infty$, gives rise to the usual expression for the power spectra

$$
\begin{equation*}
P(\omega)=\Gamma^{2} \int_{0}^{\infty} d \tau e^{i \omega \tau}\left\langle L^{\dagger}(0) L(\tau)\right\rangle \tag{29}
\end{equation*}
$$

The system correlations $\left\langle L^{\dagger}(0) L(\tau)\right\rangle$ can be computed with the quantum regression theorem.

## IV. TWO-TIME CORRELATION FUNCTION OF SYSTEM OBSERVABLES IN THE WEAK-COUPLING LIMIT

In order to evaluate the spectra from Eq. (23), it is necessary to obtain the evolution of two-time correlations of the
system observables. A theory to compute multiple time correlation functions in the non-Markovian regime has been developed in Refs. [10,11]. In this section, we briefly present some of the main results, and the formula to compute twotime correlations in the weak-coupling limit needed in Eq. (23).

The evolution equation of the quantum mean value of $A$, provided that the total initial state is $\left|\Psi_{0}\right\rangle=\left|\psi_{0}\right\rangle|0\rangle$, is equal to

$$
\begin{align*}
& \frac{d}{d t_{1}}\left\langle\Psi_{0}\right| A\left(t_{1}\right)\left|\Psi_{0}\right\rangle \\
&= i\left\langle\Psi_{0}\right|\left[H_{S}\left(t_{1}\right), A\left(t_{1}\right)\right]\left|\Psi_{0}\right\rangle \\
&+\int_{0}^{t_{1}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left[L^{\dagger}\left(t_{1}\right), A\left(t_{1}\right)\right] L(\tau)\left|\Psi_{0}\right\rangle \\
&+\int_{0}^{t_{1}} d \tau \alpha^{*}\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right| L^{\dagger}(\tau)\left[A\left(t_{1}\right), L\left(t_{1}\right)\right]\left|\Psi_{0}\right\rangle \tag{30}
\end{align*}
$$

In the same manner the evolution of the quantum mean value $\left\langle A\left(t_{1}\right) B\left(t_{2}\right)\right\rangle$ is given by

$$
\begin{align*}
& \frac{d\left\langle\Psi_{0}\right| A\left(t_{1}\right) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle}{d t_{1}} \\
& =i\left\langle\Psi_{0}\right|\left[H_{S}\left(t_{1}\right), A\left(t_{1}\right)\right] B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
& \quad+\int_{0}^{t_{1}} d \tau \alpha^{*}\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right| L^{\dagger}(\tau)\left[A\left(t_{1}\right), L\left(t_{1}\right)\right] B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
& \quad+\int_{t_{2}}^{t_{1}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left[L^{\dagger}\left(t_{1}\right), A\left(t_{1}\right)\right] L(\tau) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
& \quad+\int_{0}^{t_{2}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left[L^{\dagger}\left(t_{1}\right), A\left(t_{1}\right)\right] B\left(t_{2}\right) L(\tau)\left|\Psi_{0}\right\rangle \tag{31}
\end{align*}
$$

Equations (30) and (31) represent the evolution of quantum mean values and two-time correlations, respectively, obtained without the use of any approximation. However, it is clear that these equations are open, in the sense that quantum mean values depend on two-time correlations, while twotime correlations depend on three time correlations. In general when no approximations are made, N -time correlation depend on $N+1$-time correlations, what gives rise to a hierarchy structure of MTCF as described in Ref. [21].

For practical computations, it is necessary to make some assumptions. In our case we perform a perturbative expansion on the parameter $g$, so that the former equations become closed. Particularly, the perturbative hypothesis is applied to the operators $L(\tau)$ and $L^{\dagger}(\tau)$ that appears in Eqs. (30) and (31). The idea is to transform groups of operators with a two-time dependency on groups with a one time dependency [22]. For instance, the first dissipative term in Eqs. (30) and (31) is of the type

$$
\begin{align*}
L^{\dagger}(\tau)\{[L, A]\}\left(t_{i}\right) & =\mathcal{U}_{I}^{-1}\left(t_{i} 0\right) L\left(\tau, t_{i}\right)[L, A] \mathcal{U}_{I}\left(t_{i} 0\right) \\
& =\left\{V_{\tau-t_{i}} L^{\dagger}[L, A]\right\}\left(t_{i}\right)+O(g) \tag{32}
\end{align*}
$$

Here, we have used a Liouville superoperator $V_{t}$ which acts only on the next system operator as $V_{t} L$ $=e^{i\left(H_{S}+H_{B}\right) t} L e^{-i\left(H_{S}+H_{B}\right) t}$. In the same way, the second dissipative term in Eq. (31) can be expanded as

$$
\begin{align*}
L(\tau) B\left(t_{i+1}\right) & =\mathcal{U}_{I}^{-1}\left(t_{i+1} 0\right) L\left(\tau, t_{i+1}\right) B \mathcal{U}_{I}\left(t_{i+1} 0\right) \\
& =\left\{V_{\tau-t_{i+1}} L B\right\}\left(t_{i+1}\right)+O(g), \tag{33}
\end{align*}
$$

up to zero order. Note that just as before, the dissipative terms are already of second order due to the appearance of the second order function $\alpha(t)$, so that only zero order expansions of the rest of the quantities are needed. Then, up to second order in $g$, we obtain the following equation for quantum mean values:

$$
\begin{align*}
\frac{d}{d t_{1}}\langle & \left.\Psi_{0}\left|A\left(t_{1}\right)\right| \Psi_{0}\right\rangle \\
= & i\left\langle\Psi_{0}\right|\left\{\left[H_{S}, A\right]\right\}\left(t_{1}\right)\left|\Psi_{0}\right\rangle \\
& +\int_{0}^{t_{1}} d \tau \alpha^{*}\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left\{V_{\tau-t_{1}} L^{\dagger}[A, L]\right\}\left(t_{1}\right)\left|\Psi_{0}\right\rangle \\
& +\int_{0}^{t_{1}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left\{\left[L^{\dagger}, A\right] V_{\tau-t_{1}} L\right\}\left(t_{1}\right)\left|\Psi_{0}\right\rangle+O\left(g^{3}\right) \tag{34}
\end{align*}
$$

and for two-time correlations

$$
\begin{align*}
& \frac{d}{d t_{1}}\left\langle\Psi_{0}\right| A\left(t_{1}\right) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
&= i\left\langle\Psi_{0}\right|\left\{\left[H_{S}, A\right]\right\}\left(t_{1}\right) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
&+\int_{0}^{t_{1}} d \tau \alpha^{*}\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left\{V_{\tau-t_{1}} L^{\dagger}[A, L]\right\}\left(t_{1}\right) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
&+\int_{0}^{t_{1}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left\{\left[L^{\dagger}, A\right] V_{\tau-t_{1}} L\right\}\left(t_{1}\right) B\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
&+\int_{0}^{t_{2}} d \tau \alpha\left(t_{1}-\tau\right)\left\langle\Psi_{0}\right|\left\{\left[L^{\dagger}, A\right]\right\}\left(t_{1}\right)\left\{\left[B, V_{\tau-t_{2}} L\right]\right\}\left(t_{2}\right)\left|\Psi_{0}\right\rangle \\
&+O\left(g^{3}\right) \tag{35}
\end{align*}
$$

While the first two terms of Eq. (35) are analogous to those of Eq. (34), the equation for two-time correlations contains an additional term that does not appear in the evolution of quantum mean values. Since the quantum regression theorem (QRT) states that the evolution of two-time correlations has the same coefficients as the evolution of one-time correlations, these equations show that the theorem is not satisfied for non-Markovian interactions. The validity of the QRT have been previously discussed by some authors [23-26]. Notice that for Markovian interactions the last term of Eq. (35) vanishes, since the corresponding correlation function $\alpha\left(t_{1}-\tau\right)=\Gamma \delta\left(t_{1}-\tau\right)$ is zero in the domain of integration from 0 to $t_{2}$, so that the QRT is valid in this case. An equation
similar to Eq. (35) can be derived for initial condition $\left\langle A\left(t_{1}\right) B\left(t_{1}\right)\right\rangle$, and for $t_{2}>t_{1}$. For the Markov case, it can also be shown that this equation is equivalent to the one given by the QRT for such an initial condition. For our purposes, we shall use Eq. (35) to evaluate the evolution of nonMarkovian two-time correlations.

## V. FLUORESCENCE SPECTRA OF A TWO LEVEL ATOM

The formulas derived up to now are valid for the emission spectra of an arbitrary atom, provided that the interaction Hamiltonian has the linear form (1) and that the weak coupling approximation is valid. In this section, we show how to express the Hamiltonian for a driven two-level atom, so that a particular form for the coupling operator $L$ is given.

As in Ref. [27], we consider the interaction Hamiltonian $H_{S L}$ of the atom with the classical coherent monochromatic laser field in the usual rotating wave approximation form [3]

$$
\begin{equation*}
H_{S L}=\epsilon\left(\sigma_{21} e^{-i\left(\omega_{L} t+\phi_{T}\right)}+\sigma_{12} e^{i\left(\omega_{L} t+\phi_{T}\right)}\right), \tag{36}
\end{equation*}
$$

where $\omega_{L}$ is the frequency of the laser in the coherent state $\alpha e^{-i \omega_{L} t}$, with $\alpha=|\alpha| e^{-i \phi_{L}}$. The quantity $\epsilon=\mathbf{d}_{21} \epsilon$ is the Rabi frequency, and $|\epsilon|=\sqrt{\omega_{L} / 2 \epsilon_{0} v} \sqrt{N_{L}} e_{L}$ the laser field magnitude, where $v$ is the volume of the cavity, and $\bar{N}_{L}=\left|\alpha^{2}\right|$ and $e_{L}$, are, respectively, the mean number of photons and the polarization of the laser mode. The phase $\phi_{T}$ is a global phase defined as $\phi_{T}=\phi_{L}-\pi / 2$. Because of the magnitude of the laser field, the Hamiltonian $H_{S L}$ should be considered as part of the noninteracting Hamiltonian $H_{0}$, where in this case $H_{0}=H_{S}+H_{B}+H_{S L}$. We can eliminate the explicit dependence on the laser frequency by changing to a rotating frame with a frequency $\omega_{L}$ by using the unitary operator

$$
\begin{equation*}
U_{t}=e^{\left[i \omega_{L} t+i \phi_{T}\right]\left[\Sigma_{\lambda} a_{\lambda}^{\dagger} a_{\lambda}+\left(\sigma_{22}-\sigma_{11}\right)\right]} \tag{37}
\end{equation*}
$$

If $\Delta_{\lambda}=\omega_{\lambda}-\omega_{L}$ and $\Delta_{S L}=\omega_{S}-\omega_{L}$ the rotated Hamiltonian $H^{\prime}$ $=H_{0}^{\prime}+H_{I}^{\prime}$, where

$$
\begin{equation*}
H_{0}^{\prime}=\sum_{\lambda} \Delta_{\lambda} a_{\lambda}^{\dagger} a_{\lambda}+\frac{1}{2} \omega_{S} \sigma_{3}+\epsilon\left[\sigma_{21}+\sigma_{12}\right] \tag{38}
\end{equation*}
$$

and

$$
\begin{equation*}
H_{I}^{\prime}=i \sum_{\lambda} g_{\lambda}\left(\sigma_{12} a_{\lambda}^{\dagger}-a_{\lambda} \sigma_{21}\right) \tag{39}
\end{equation*}
$$

can still be expressed in a simpler way by projecting it into the dressed atomic basis. The new Hamiltonian $\widetilde{H}=V^{-1} H^{\prime} V$, where

$$
V=\left(\begin{array}{cc}
\mathbf{c} & -\mathbf{s}  \tag{40}\\
\mathbf{s} & \mathbf{c}
\end{array}\right)
$$

will be of the form (1). The constants appearing in the unitary transformation matrix $V$ are $\mathbf{c}=\cos \phi$ and $\mathbf{s}=\sin \phi$, where the angle $\phi$ is given by $\sin ^{2} \phi=\frac{1}{2}[1$ $-\operatorname{sgn}\left(\Delta_{S L}\right) / \sqrt{\left.\epsilon^{2} / \Delta_{S L}^{2}\right)+1}$. The noninteracting dressed state Hamiltonian $\widetilde{H}_{0}=\widetilde{H}_{S}+\widetilde{H}_{B}+\widetilde{H}_{S L}$ is equal to

$$
\begin{equation*}
\tilde{H}_{0}=\Omega R_{3}+\sum_{\lambda} \Delta_{\lambda} a_{\lambda}^{\dagger} a_{\lambda} \tag{41}
\end{equation*}
$$

and the interaction Hamiltonian $\widetilde{H}_{I}$ has the same form as in Eq. (1) once the interaction operator $L$ is defined as

$$
\begin{equation*}
L=\mathbf{c} \cdot \mathbf{s} R_{3}+\mathbf{c}^{2} R_{12}-\mathbf{s}^{2} R_{21} \tag{42}
\end{equation*}
$$

Here, $R_{i j}=|\tilde{i}\rangle\langle\tilde{j}|$ are the atomic operators defined in the dressed state basis $\{|\widetilde{1}\rangle,|\widetilde{2}\rangle\}$, and $R_{3}=R_{22}-R_{11}$. The quantity $\Omega=\left[\epsilon^{2}+\Delta_{S L}^{2} / 4\right]^{1 / 2}$ is called the generalized Rabi frequency.

According to Eq. (40), the relation between the dressed atomic operators and the bare atomic operators is the following:

$$
\begin{gather*}
\sigma_{12}=\mathbf{c} \cdot \mathbf{s} R_{3}+\mathbf{c}^{2} R_{12}-\mathbf{s}^{2} R_{21}  \tag{43}\\
\sigma_{21}=\mathbf{c} \cdot \mathbf{s} R_{3}-\mathbf{s}^{2} R_{12}+\mathbf{c}^{2} R_{21} \\
\sigma_{3}=\left(\mathbf{c}^{2}-\mathbf{s}^{2}\right) R_{3}-2 \mathbf{c} \cdot \mathbf{s}\left(R_{12}+R_{21}\right)
\end{gather*}
$$

It can be easily verified that the case of spontaneous emission corresponds to switching off the laser, and therefore $L$ $=\sigma_{12}$.

Once the coupling operator $L$ is known, the two time correlation $\left\langle L^{\dagger}(\tau) L\left(\tau^{\prime}\right)\right\rangle$ can be easily computed with Eq. (35). Now, in order to obtain the emission spectra with Eq. (23), it is necessary to derive the Laplace transform of the spatial dependent correlation function $\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)$. Notice that the interaction with the field enters into formula (23) in two ways. Indirectly, through the correlation function $\alpha\left(t_{1}-\tau\right)$ entering in Eq. (35). Directly, through the function $\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)$.

Let us make here a remark. In order to compute the atomic dynamical equations, we assume that the field is composed of extended modes, in the sense that they correspond to a real wave vector. Hence, the spatial dependency is assumed to be $\exp (i \mathbf{k} \cdot \mathbf{r}) \approx 1$ within the optical region. This assumption corresponds to consider the photonic crystal as an infinite and perfect structure, so that evanescent modes are, in principle, not present. Nevertheless, a defect placed in the crystal (in the our case an emitting atom) may substain such modes: once an impurity atom with resonant frequency within the gap region is placed within the crystal, the field that it emits is in the form of evanescent modes (i.e., with imaginary wave vector). Then, the approximation $\exp (i \mathbf{k} \cdot \mathbf{r}) \approx 1$ is no longer valid to describe the emitted field. As will be explained in detail later, this fact is very relevant to compute the emission spectra, which appears to be strongly dependent on the atom-detector distance through the spatially dependent function $\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)$.

In the next section, we present the calculus of $\alpha(t-\tau)$ and $\mathcal{S}^{*}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)$ necessary to compute the system fluctuations and the fluorescence spectra, respectively. Both quantities depend on the particular interaction considered. We study them for the case of an atom interacting with the radiation field within a photonic crystal.

## VI. ENVIRONMENT CORRELATION FUNCTIONS FOR AN ATOM IN A PBG

In this section, we calculate the correlation function characteristic of the dipolar coupling of a two level atom to the modified radiation field within a PC. In that case, the coupling constants entering in Eq. (2) are

$$
\begin{equation*}
g_{\lambda}=-i \sqrt{\frac{1}{2 \epsilon_{0} \omega_{\lambda}}} \omega_{12} \hat{\mathbf{e}}_{\mathbf{k} \sigma} \cdot \mathbf{d}_{12} e^{i \mathbf{k} \cdot \mathbf{r}} \tag{44}
\end{equation*}
$$

where $\mathbf{d}_{12}$ is the atom dipolar moment, $\hat{\mathbf{e}}_{\mathbf{k} \sigma}$ is the unit vector in the direction of the polarization $\sigma$ for a given wave vector $\mathbf{k}$, and $\epsilon_{0}$ is the electric vacuum permittivity. In the last expression we have already expanded the potential vector in terms of plane waves. This is also valid for radiation fields in photonic crystals, but with the additional requirement that these plane waves are periodic, i.e., satisfy the Bloch theorem. As we have already noted in the introduction, the presence of a periodic dielectric structure produces the appearance of optical bands, which are themselves periodic structures in the reciprocal lattice. Periodic eigenvalues (the frequencies), correspond to periodic eigenvectors (the field). However, although the quantity $\lambda \equiv\{\mathbf{k}, \sigma, n\}$, where $n$ denotes each band, this index will be here dropped, since we consider the dynamics of the two level system with rotating frequency nearby (or inside) a single band.

In order to calculate $\alpha(t)$, and $\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t\right)$, the sums appearing in Eqs. (2) and (20), respectively, have to be performed, by using the functions $g_{\lambda}$ as defined in Eq. (44). In that way we have

$$
\begin{equation*}
\alpha(\tau)=\gamma\left(\frac{a}{2 \pi}\right)^{3} \sum_{\sigma} \int_{1 \mathrm{BZ}} d \mathbf{k} \frac{\left|\hat{e}_{\mathbf{k}, \boldsymbol{\sigma}} \cdot \hat{u}_{d}\right|^{2}}{\omega(\mathbf{k})} e^{-i \omega(\mathbf{k}) \tau} \tag{45}
\end{equation*}
$$

and

$$
\begin{align*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \tau\right)= & \gamma\left(\frac{a}{2 \pi}\right)^{3} \sum_{\sigma} \int_{1 \mathrm{BZ}} d \mathbf{k} \frac{\left|\hat{e}_{\mathbf{k}, \sigma} \cdot \hat{u}_{d} \| \hat{e}_{\mathbf{k}, \sigma} \cdot \hat{u}_{d}^{\mathcal{D}}\right|}{\omega(\mathbf{k})} e^{-i \omega(\mathbf{k}) \tau} \\
& \times e^{i\left(\mathbf{r}-\mathbf{r}_{a}\right) \mathbf{k}} \tag{46}
\end{align*}
$$

where $\hat{u}_{d}^{\mathcal{D}}$ is the unitary vector corresponding to the dipolar moment of the detector, and the integrals are performed over the first Brillouin zones of the crystal. Also, it is clear from Eq. (44) that the constant $\gamma=\omega_{12}^{2} d_{12}^{2} /\left(2 \epsilon_{0}\right)$. In order to calculate the integrals appearing in Eqs. (45) and (46), it is clear that the dispersion relation $\omega(\mathbf{k})$ is needed.

The dynamics near the edge of the band are often described through an effective mass approximation to the full dispersion relation, based on its expansion in the vicinity of the band edge (see, for instance, Refs. [27-30]). Within this approximation, and considering the simplest case of a cubic lattice, the dispersion relation has the parabolic form $\omega(\mathbf{k})$ $=\omega_{c}+\mathcal{A}\left(\mathbf{k}-\mathbf{k}_{0}\right)^{2}$, where $\mathbf{k}_{0}$ is the origin of the first Brillouin zone of the crystal (which is the unitary cell in $\mathbf{k}$ space) about which we perform the expansion in each direction, $\omega_{c}$ is the frequency of the band edge, and $\mathcal{A}$ is a constant that depends on the specific photonic crystal considered. Although it is valid for the description of several important physical phenomena occurring in PBG materials
[2,27,31-33], this model of dispersion relation becomes inaccurate for a good description of the short time dynamics. The dispersion relation proposed in Ref. [14] overcomes this problem by reproducing the periodic structure of the bands

$$
\begin{equation*}
\omega(\mathbf{k})=A+\frac{B}{3}\left[\cos \left(k_{x} a\right)+\cos \left(k_{y} a\right)+\cos \left(k_{z} a\right)\right] . \tag{47}
\end{equation*}
$$

Notice that this function reduces to the parabolic model for $\mathbf{k} \approx \mathbf{k}_{0}$. For a three dimensional band structure with the periodic dispersion relation (47), the correlation function can be obtained from the Eq. (45) once it has been assumed that the function $\left|\hat{e}_{k, \sigma} \cdot \hat{u}_{d}\right|^{2} / \omega(\mathbf{k})$ is a slowly varying function in the 1 BZ . The resulting integral is analytical, and gives rise to the following result:

$$
\begin{equation*}
\alpha_{3 \mathrm{D}}(\tau) \cong \gamma \sum_{\sigma} \frac{\left|\hat{e}_{\mathbf{k}_{0}, \sigma} \cdot \hat{u}_{d}\right|^{2}}{\omega\left(\mathbf{k}_{0}\right)} e^{-i A \tau} J_{0}^{3}\left(\frac{B \tau}{3}\right) \cong g_{3 \mathrm{D}}^{2} e^{-i A \tau} J_{0}^{3}\left(\frac{B \tau}{3}\right), \tag{48}
\end{equation*}
$$

where $J_{0}$ is the zero order Bessel function.
Let us now consider the spatial dependent correlation function $\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, t, \tau\right)$. As before, we assume that the quantity $\frac{\left|\hat{\mathbf{e}}_{\mathbf{k}, \sigma} \cdot \hat{u}_{d} \| \hat{\hat{N}}_{\mathbf{k}, \sigma} \cdot \hat{\sigma}_{\vec{d}}\right|}{\omega(\mathbf{k} \mid}$ is a slowly varying function in the reciprocal space (particularly in the region nearby the symmetric point $\mathbf{k}_{0}$ ), in therefore we can express the function (46) as

$$
\begin{align*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \tau\right)= & \gamma\left(\frac{a}{2 \pi}\right)^{3} \sum_{\sigma} \frac{\left|\hat{e}_{\mathbf{e}_{0}, \sigma} \cdot \hat{u}_{d} \| \hat{e}_{\mathbf{k}_{0}, \sigma} \cdot \hat{u}_{d}^{\mathcal{D}}\right|}{\omega\left(\mathbf{k}_{0}\right)} \int_{1 \mathrm{BZ}} d \mathbf{k} e^{-i \omega(\mathbf{k}) \tau} \\
& \times e^{i\left(\mathbf{r}-\mathbf{r}_{a}\right) \cdot \mathbf{k}} . \tag{49}
\end{align*}
$$

Its Laplace transform is then given by

$$
\begin{align*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)= & -i \gamma\left(\frac{a}{2 \pi}\right)^{3} \sum_{\sigma} \frac{\left|\hat{e}_{\mathbf{k}_{0}, \sigma} \cdot \hat{u}_{d} \| \hat{e}_{\mathbf{k}_{0}, \sigma} \cdot \hat{u}_{d}^{\mathcal{D}}\right|}{\omega\left(\mathbf{k}_{0}\right)} \int_{1 \mathrm{BZ}} d \mathbf{k} \\
& \times \frac{e^{i\left(\mathbf{r}-\mathbf{r}_{a}\right) \cdot \mathbf{k}}}{(\omega(\mathbf{k})-\omega)} . \tag{50}
\end{align*}
$$

In order to perform the last integrals analytically, it is convenient to consider the parabolic dispersion relation and the limit in which the detector and the emitting atom are far away from each other.

In that way, we can make a change of variable $\mathbf{q}=\mathbf{k}-\mathbf{k}_{0}$, so that we have $\omega(\mathbf{k})=\omega_{c}+\mathcal{A} \mathbf{q}^{2}$ and extend the limits of integration to infinity. Then

$$
\begin{align*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)= & -i Q \int d^{3} q \frac{e^{i \mathbf{q} \cdot \mathbf{d}}}{\omega(q)-\omega}= \\
& -i Q \frac{2 \pi}{d} \int_{0}^{\infty} d q q \frac{e^{i k d}-e^{-i k d}}{\omega(q)-\omega} \tag{51}
\end{align*}
$$

We have also assumed that $\mathbf{q}=q(\sin \theta \cos \phi, \sin \theta \sin \phi$, $\cos \theta$ ), and the vector $\mathbf{d}=\mathbf{r}-\mathbf{r}_{a}$ being parallel to the $z$ axis. The constant


FIG. 1. (Color online) Solid line: Localization length $\xi$ with respect to frequency showing a decreasing for increasing detunings from the band edge. Dotted (dashed) lines: Real (imaginary) parts of the function (55).

$$
\begin{equation*}
Q=\gamma\left(\frac{a}{2 \pi}\right)^{3} e^{i \mathbf{k}_{0} \cdot \mathbf{d}}\left(1-\frac{\left(\mathbf{k}_{0} \cdot \mathbf{u}_{d}\right)^{2}}{k_{0}^{2}}\right)\left(1-\frac{\left(\mathbf{k}_{0} \cdot \mathbf{u}_{d}^{\mathcal{D}}\right)^{2}}{k_{0}^{2}}\right), \tag{52}
\end{equation*}
$$

provided that the $d_{12}^{\mathcal{D}}=d_{12}$. Considering that $\theta$ and $\theta^{\mathcal{D}}$ are the angles between $\mathbf{k}_{0}$ and $\mathbf{u}_{d}$ and $\mathbf{u}_{d}^{\mathcal{D}}$, respectively, we have simply

$$
\begin{equation*}
Q=\gamma\left(\frac{a}{2 \pi}\right)^{3} e^{i \mathbf{k}_{0} \cdot \mathbf{d}} \sin ^{2} \theta \sin ^{2} \theta^{\mathcal{D}} \tag{53}
\end{equation*}
$$

Notice that in our case, since we are dealing with a cubic lattice, only one symmetric point $\mathbf{k}_{0}$ and the dispersion relation around it is being considered. In a more general case, the band edge is associated with a finite collection of symmetryrelated point $\mathbf{k}_{0}^{i}$, so that the calculus has to be taken into account the dispersion relation near each of the points $\omega(\mathbf{k})=\omega_{c}+\mathcal{A}\left(\mathbf{k}-\mathbf{k}_{0}^{i}\right)^{2}$. In such a case, the constant

$$
\begin{equation*}
Q=\gamma\left(\frac{a}{2 \pi}\right)^{3} \sum_{i} e^{i \mathbf{k}_{0}^{i} \cdot \mathbf{d}} \sin ^{2} \theta_{i} \sin ^{2} \theta_{i}^{\mathcal{D}} \tag{54}
\end{equation*}
$$

Performing the last integral of Eq. (51), we just obtain

$$
\begin{equation*}
\mathcal{S}\left(\mathbf{r}, \mathbf{r}_{a}, \omega\right)=-\frac{i Q 2 \pi^{2}}{d \mathcal{A}} e^{-d / \xi} \tag{55}
\end{equation*}
$$

where $\xi=1 / \sqrt{\frac{\omega_{c}-\omega}{\mathcal{A}}}[34]$.
As displayed in Fig. 1, an increasing distance in frequencies with respect to the band edge, gives rise to a decreasing localization length $\xi$, and therefore to a decreasing value of the function $\mathcal{S}$ [35]. In all the numerical results we consider units of $\omega_{L}=1, \mathcal{A}=1$ (which implies that distances are in units of $\left.\sqrt{\frac{\mathcal{A}}{\omega_{L}}}\right)$, and the frequency origin in $\omega_{C}$.

Numerical results. In this section we compute the florescence spectra of a two-level atom placed as an impurity in a


FIG. 2. (Color online) A comparison between the Mollow triplet obtained with Markovian interaction (solid line) and the triplet obtained with the non-Markovian interaction (dotted line) within a PC is here shown for $\Omega=0.2$ (top) and $\Omega=0.3$ (lower). The tree components of the triplet are placed within the band, that corresponds to frequencies $A-B<\omega<A+B$, where $A=1$ and $B=1$ (units of $\omega_{L}$ $=1)$. The asymmetry between the left and the right side-bands described in Ref. [36] is also described with our model.
photonic crystal. To this order, we will study two cases, the florescence spectra when no spatial dependency is considered, Eq. (27), and when the spatial dependency is considered (i.e., the distance from the emitting atom to the detector), in Eq. (23). In both cases, we need to compute the Laplace transform of the system fluctuations $\left\langle L^{\dagger}(\tau) L\left(\tau^{\prime}\right)\right\rangle$, obtained with Eq. (35) by using the coupling operator $L$ given by Eq. (42). We show that the spatial dependency

a)
b)

FIG. 3. (Color online) Schematic view of the situations we consider in our numerical calculations. (a) and (b) correspond to the two cases analyzed when no spatial dependency is considered. (c) is the setup considered when studying the spatial dependency.


FIG. 4. (Color online) In contrast with Fig. 2, now we place the left and right emission lines within the gap. When the distance between atom and detector is not considered, there is no filtering of the fluorescent emission that occurs within the gap.
should be considered when part of the radiation is emitted in the form of evanescent waves, since in that case the approximation $\exp (i \mathbf{k} \cdot \mathbf{r}) \approx 1$ is no longer appropriate. Evanescent modes are emitted for frequencies within the gap.

No spatial dependency considered. We analyze here the atomic fluorescence spectra according to Eq. (27), by using the correlation function (48). Figure 2 shows a case in which we have chosen the Rabi and the laser frequencies $(\Omega$ and $\left.\omega_{L}\right)$ in such a way that the we have placed the three components of the Mollow triplet inside the band. This corresponds to the setup of Fig. 3(a). The appearance of a three component fluorescence spectra for a driven atom is due to the dressing of the atomic levels by the laser [3]. In Fig. 4 we show a case in which the two side-bands of the triplet are placed inside the gap [see Fig. 3(b)]. As we will clearly show in the next section, this two lines corresponds to emission of evanescent modes, which have purely imaginary wave vectors. Hence, the approximation $\exp (i \mathbf{k} \cdot \mathbf{r}) \approx 1$ over estimates their magnitude, and a more accurate description of the spatial dependency is needed. Only with such description we will be able to account for the fact that evanescent modes should not be detected in the far field. This is done in the next section.

Spatial dependency considered. From the last section we have concluded that for a correct description of the emission from the gap, it is very important to take into account the spatial dependency, i.e., the distance of the detector from the emitting atom. To this order, we consider the Eq. (23), and the Laplace transform of the spatial correlation function (55) that is obtained when considering the parabolic dispersion relation $\omega(\mathbf{k})=\omega_{c}+\mathcal{A}\left(\mathbf{k}-\mathbf{k}_{0}\right)^{2}$. Moreover, in order to compute the system fluctuations (35), we take the (not spatially dependent) correlation function corresponding to this dispersion relation, which according to Ref. [27] is


FIG. 5. (Color online) Power spectra $P(\omega)$ for different values of the Rabi frequency $\Omega$. Solid line: $\Omega=0.55$, dotted line: $\Omega=0.6$, and dashed line: $\Omega=0.7$. The distance of the detector is $d=10$, small enough as to detect evanescent modes coming from the left sideband of the Mollow triplet, which is placed in the gap.

$$
\begin{equation*}
\alpha(\tau) \cong \beta^{1 / 2} \frac{e^{i\left[\pi / 4-\omega_{c} \tau\right]}}{\tau^{3 / 2}} \tag{56}
\end{equation*}
$$

This correlation is just Eq. (48) taking the long time limit, so that the constants can be related as [14]

$$
\begin{equation*}
\beta^{1 / 2}=\frac{g_{3 \mathrm{D}}^{2}}{8}\left(\frac{6}{B}\right)^{3 / 2} \tag{57}
\end{equation*}
$$

When using the parabolic dispersion relation, only two frequency regions have to be taken into account: a gap, for $\omega$


FIG. 6. (Color online) Power spectra $P(\omega)$ times $d$ for different values of the the detector distance $d$. Solid line: $d=0.1$, dotted line: $d=10$, and dashed line: $d=100$. The Rabi frequency is $\Omega=0.55$, so that the left sideband of the Mollow triplet is emitted in gap frequencies. For large distances, the left sideband is no longer detected, and a complete filtering of the spectra within the gap is obtained.
$<\omega_{c}$, and a band for $\omega>\omega_{c}$. We chose our parameters $\left(\Delta_{\mathrm{AL}}=0\right.$ and $\left.\Omega\right)$ such that the left Mollow component is placed in the gap, while central and right component have frequencies within the band [see Figs. 3(c) for a schematic view]. In Fig. 5 we show the fluorescence spectra for a certain distance $d$ of the detector which is small enough as to still detect some evanescent modes from the gap. Different choices of $\Omega$ shows that the more deep in the gap the emission occurs, the less we are efficient to detect it, due to the decreasing of the localization length. Figure 6 shows the spectra for a fixed value of $\Omega$, but varying the distance of the detector, in such a way that for $d$ large enough, a complete suppression of fluorescence from the gap (corresponding to the left side-band) occurs.

Concerning this, two important remarks have to be made here. We have seen that there is emission within the gap in the form of evanescent modes. This result cannot be recovered with the Markovian approximation, when using emission rates computed with the Fermi golden rule (see, for instance, Ref. [37]). In addition, contrary to Refs. [38] and [36], where a filter $\Theta\left(\omega-\omega_{c}\right)$ is introduced for emission frequencies within the gap, in the description presented here no filtering effect occurs unless the spatial dependency is considered, and the detector is placed at a distance larger than the localization length of the field. A second remark concerns the difference between the spontaneous and the fluorescence spectra. In the spontaneous emission spectra, some radiation is detected within the gap region even when placing the detector far away from the emitting atom. The reason is that even for frequencies within the gap there is a transient field, the so-called diffusive field, that is emitted at the beginning of the atomic evolution. This field is not evanescent but propagates, and therefore carries away some energy from the atom. Hence, it is responsible of the slight dissipation that occurs at the initial time steps of the evolution [14,27]. A description of this field including its spatial dependency is made in Ref. [39], by solving the dynamics in the one photon sector [40].

In contrast, the fluorescence spectra deals with the signal emitted by the atom in the stationary limit, and therefore the transient field is no longer detected. For that reason, the radiation within the gap is only composed of evanescent modes, and cannot be observed when placing the detector far away from the atom.

In summary, it is true that the filtering appears as a consequence of the special form of the dispersion relation of the field in the PC. However, we show here that this filtering only occurs provided that the signal is in the stationary limit, and the distance between the emitting source and the detector is large.

## VII. CONCLUSIONS

We present in this paper a general formula to compute the emission spectra of an atom linearly coupled to an environment with non-Markovian interaction, and within the weak coupling approximation. This formula is explicitly dependent on the distance between the emitting atom and the detector $d$. We apply it to compute the fluorescence spectra of a twolevel atom in a photonic crystal. In those materials, the electromagnetic field have very specific features, showing a gap or range of frequencies where the propagating modes are not allowed. In that frequency region, the emission spectra depends critically on $d$, since the emission occurs in the form of evanescent (nonpropagating) modes, which decay exponentially from the atom. For that reason, we show that in the far field these modes are not detected. We note that the description of this evanescent field can only be made with a non-Markovian description, since a Markov theory, with dissipation rates calculated through the Fermi golden rule, predicts no emission at all within the gap region.

## ACKNOWLEDGMENTS

We thank H. Carmichael, G. C. Hegerfeldt, A. Ruíz, and L. S. Schulman for their comments and G. Nicolis. P. Gaspard, J. I. Cirac, and W. T. Strunz for support and encouragement. We thank R. F. O'Connell for pointing out some important references concerning the Quantum Regression Theorem. This work has been supported by the Gobierno de Canarias (Spain) (Grant No. PI2004/025) and Ministerio de Ciencia y Tecnología of Spain (Grants No. FIS2004-05687 an No. FIS2007-64018). I. de Vega is financially supported by a Ministerio de Ciencia y Tecnología (Grant No. EX20060295).
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