

Statistical Physics II.
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Degenerate Fermi Gas
White dwarfs

- In this section we shall study a simple model of White Dwarfs.

- Some numbers

```
<< PhysicalConstants`  
<< Units`
```

```
StarDensity = 107 Gram / Centimeter3
```

```
10 000 000 Gram  
-----  
Centimeter3
```

```
Sm = CGS [1.9891 × 1030 Kilogram]
```

```
1.9891 × 1033 Gram
```

```
Tc = 107 Kelvin
```

```
10 000 000 Kelvin
```

```
kB = BoltzmannConstant
```

```
1.38065 × 10-23 Joule  
-----  
Kelvin
```

- The thermal energy E_T corresponding to the temperature T_c

```
ET = Convert [kB Tc, ElectronVolt]
```

```
861.734 ElectronVolt
```

- This energy is much higher than the Helium ionization energy

```
ElementData[2, "IonizationEnergies", "Units"]
Io = ElementData[2, "IonizationEnergies"]
```

KilojoulesPerMole

{2372.3, 5250.5}

```
IO1 = SI [Io 1000 Joule / Mole]
```

{ 3.9393×10^{-18} Joule, 8.71867×10^{-18} Joule}

```
Convert [IO1 [[1]], ElectronVolt]
Convert [IO1 [[2]], ElectronVolt]
```

24.5872 ElectronVolt

54.4176 ElectronVolt

- Now you can compare the to energies!

■ The Fermi Energy

- Helium Mass

```
ElementData [He, "Mass"]
```

```
HeMass = 4 / AvogadroConstant Gram / Mole
```

6.64216×10^{-24} Gram

```
ElectronDensity = SI [2 StarDensity / HeMass]
```

$\frac{3.01107 \times 10^{36}}{\text{Meter}^3}$

$\hbar = \text{PlanckConstantReduced}$

$m_e = \text{ElectronMass}$

$$\epsilon_F = \frac{\hbar^2}{2 m_e} (3 \pi^2 \text{ElectronDensity})^{2/3} / .$$

$$\left\{ \text{Joule} \rightarrow \text{Kilogram Meter}^2 / \text{Second}^2, \left(\frac{1}{\text{Meter}^3} \right)^{2/3} \rightarrow 1 / \text{Meter}^2 \right\}$$

$T_F = \% / k_B / . \left\{ \text{Joule} \rightarrow \text{Kilogram Meter}^2 / \text{Second}^2 \right\}$

`Convert[ϵ_F , ElectronVolt]`

$$\frac{1.21823 \times 10^{-13} \text{ Kilogram Meter}^2}{\text{Second}^2}$$

$$8.82357 \times 10^9 \text{ Kelvin}$$

$$760357. \text{ ElectronVolt}$$

- We see that T_F / T_c is such that we can consider that the electron Fermi gas is in its ground state

T_F / T_c

$$882.357$$

- Estimation of Chandrasekhar's critical Mass

`h = PlanckConstantReduced`

`g = GravitationalConstant`

`mp = ProtonMass`

`c = SpeedOfLight`

`Sm = SI[SolarMass]`

`factor = $\frac{8 \text{ mp}}{9 \pi}$;`

$$M_0 = \left(\frac{27 \pi}{64 \alpha} \right)^{3/2} \left(\frac{h c}{g \text{ mp}^2} \right)^{3/2} /. \{ \text{Newton} \rightarrow \text{Joule} / \text{Meter} \} /. \\ \{ \text{Joule} \rightarrow \text{Kilogram Meter}^2 / \text{Second}^2 \}$$

`1.05457 × 10-34 Joule Second`

`$\frac{6.67428 \times 10^{-11} \text{ Meter}^2 \text{ Newton}}{\text{Kilogram}^2}$`

`1.67262 × 10-27 Kilogram`

`$\frac{299792458 \text{ Meter}}{\text{Second}}$`

`1.9891 × 1030 Kilogram`

`$3.3616 \times 10^{57} \left(\frac{1}{\alpha} \right)^{3/2}$`

`$M_0 \text{ factor} / Sm /. \{ \alpha \rightarrow 1 \}$`

`$M_0 \text{ factor} / Sm /. \{ \alpha \rightarrow 2 / 3 \}$`

`$M_0 \text{ factor} / Sm /. \{ \alpha \rightarrow 1.1 \}$`

`0.799806`

`1.46934`

`0.693259`

`SolarMass`

`SolarMass`